

Dawn - Overview, Science Objectives, Mission Progress

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For PI Chris Russell



Presentation to Decadal Survey
Primitive Bodies Panel,
Washington, DC, Sep 2009

Spacecraft configuration, assembly for launch



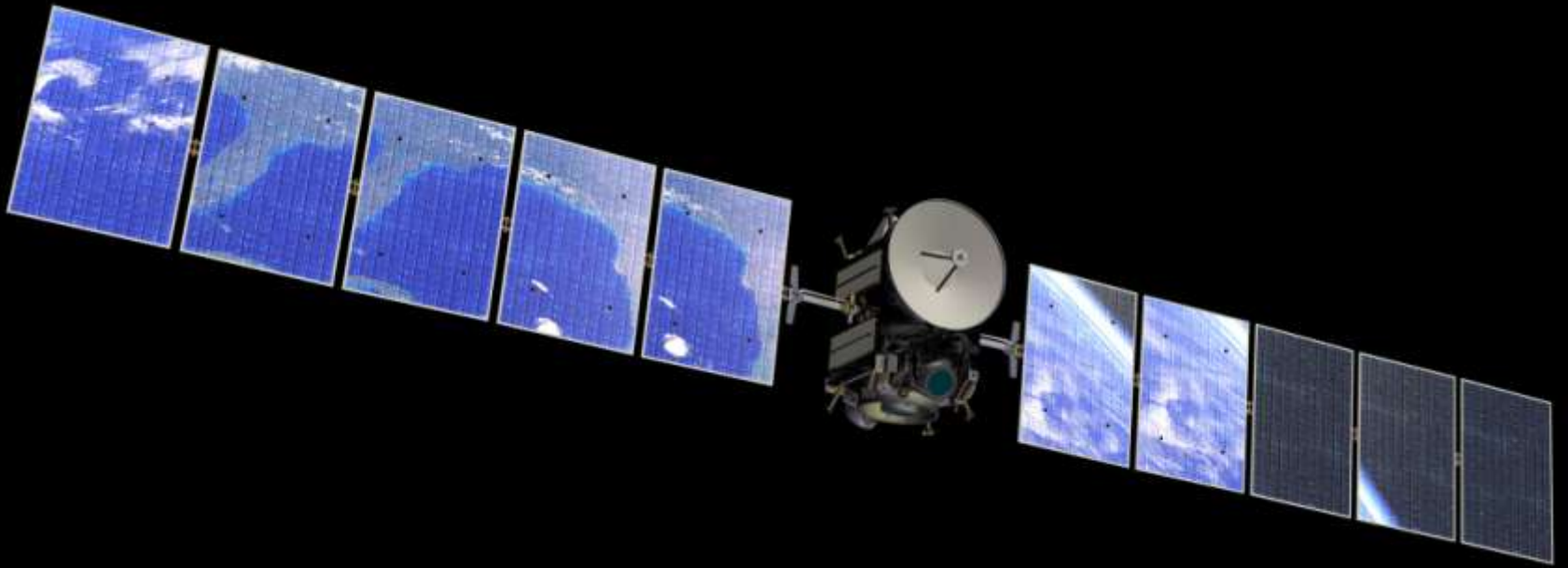
Dawn spacecraft
being lifted to the
top of its Delta II
launch vehicle



Dawn launch September 2007



Artist's view of Dawn spacecraft

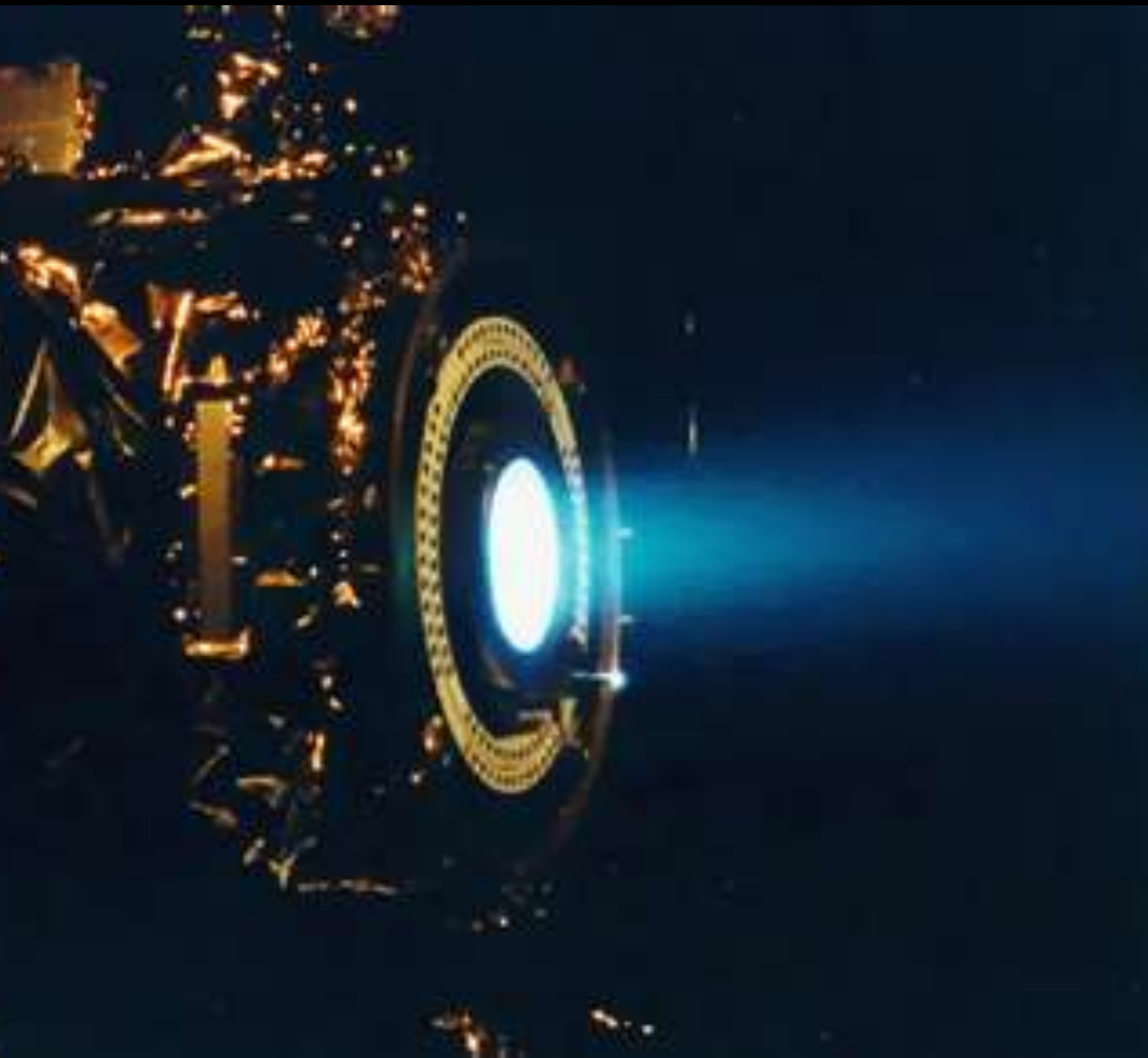


Solar arrays deployed



19.7 meter wingspan

Ion propulsion system



Dawn mission is only possible within Discovery cost cap using IPS

Dawn will have 2.5X ΔV ever obtained by a spacecraft with its propulsion system

Dawn will thrust for >5 years in its 8-year mission, 3X the longest powered flight

Dawn objectives at Vesta

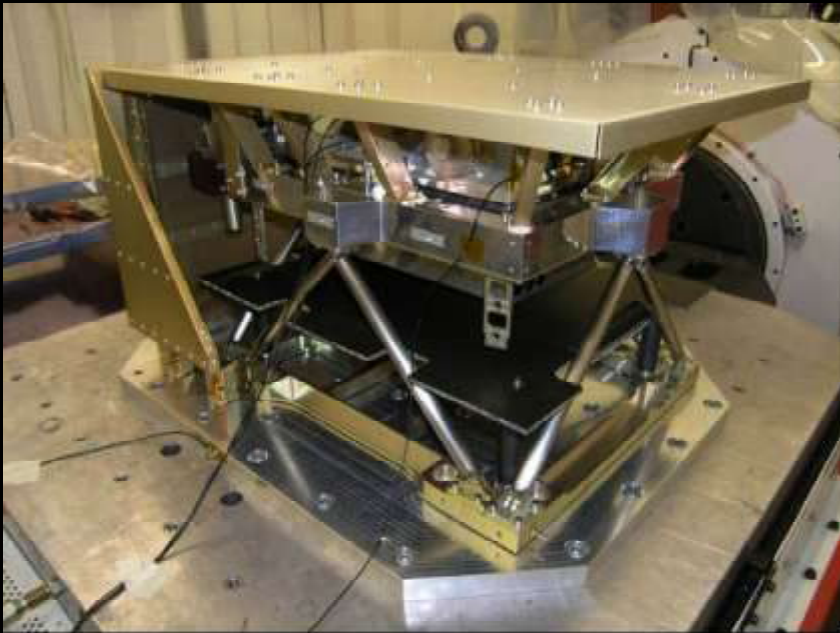
- Map the surface of Vesta in eight filters with 100-m spatial resolution to examine its geology and cratering history



Framing Camera (FC),
contributed by Germany
(Max-Planck Institute and
German Aerospace Center)

Dawn objectives at Vesta

- Determine the mineral composition of the surface at ~800-m resolution, and selected areas at 200-m or better
- Perform detailed photometric and spectroscopic examination of areas of interest, such as the walls of the southern crater

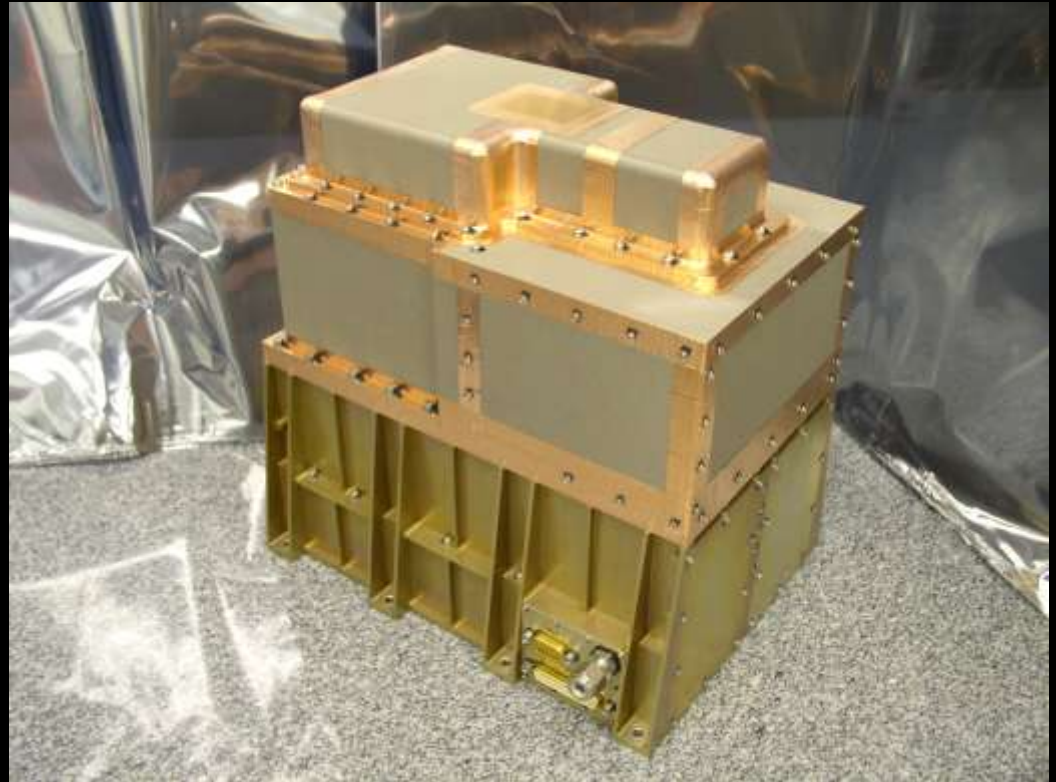


Visible and Infrared Spectrometer (VIR), contributed by Italy (Agenzia Spaziale Italiana)

Dawn objectives at Vesta

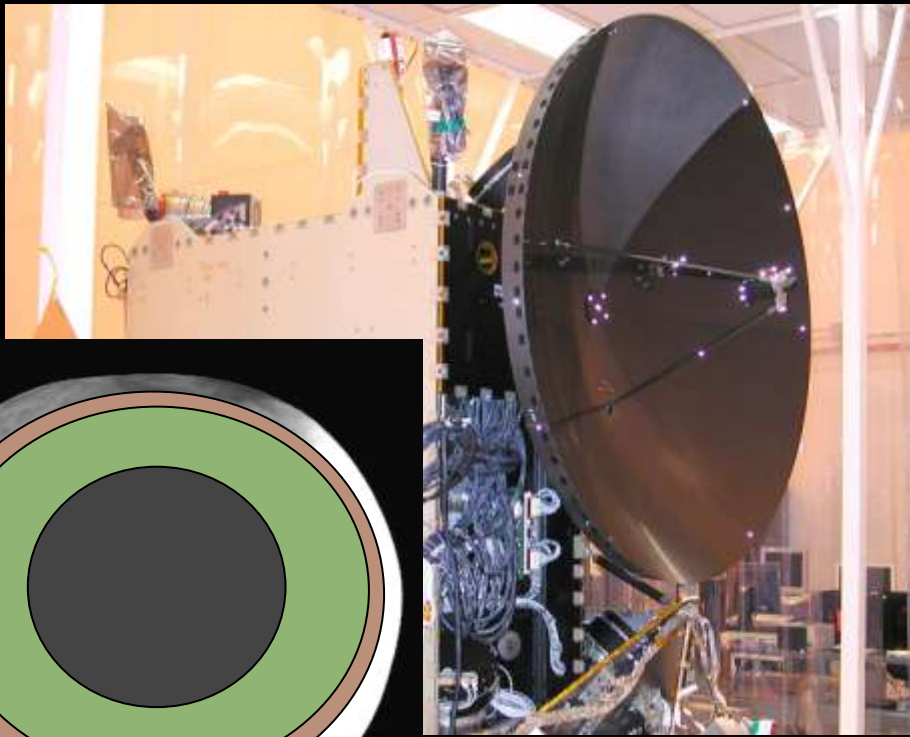
- Map elemental abundances to ~ 300 -km resolution or better

Gamma Ray and Neutron Detector (GRaND), provided by USA (Los Alamos National Laboratory)



Dawn objectives at Vesta

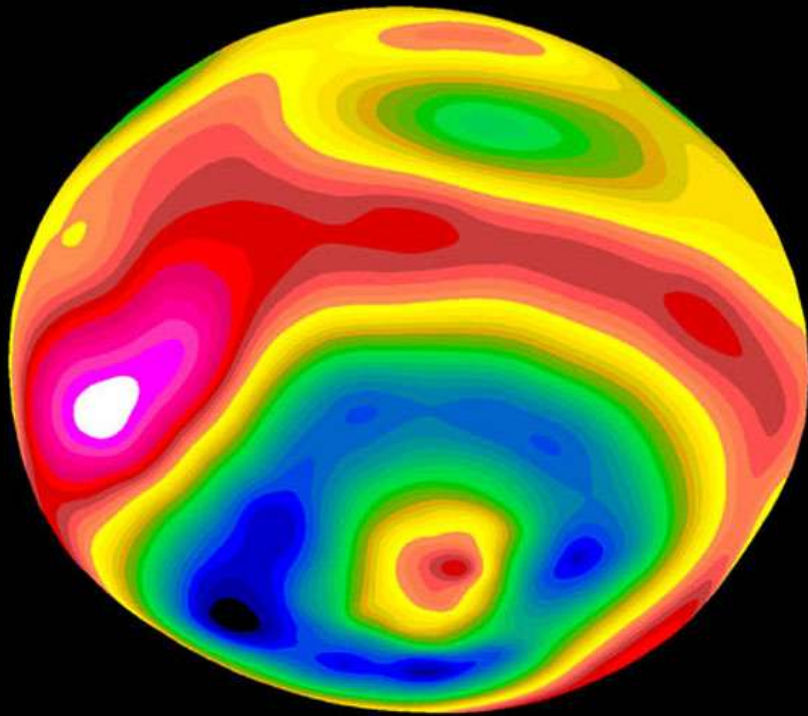
- Map the gravity field to 90-km half-wavelength resolution, to characterize crust and mantle density variations, and possibly detect a core



Gravity science using spacecraft orbit tracking via radiometric navigation, provided by USA (Orbital Sciences Corporation and JPL)

Dawn objectives at Vesta

- Resolve the topography of Vesta to 10-m height at 100-m spatial resolution



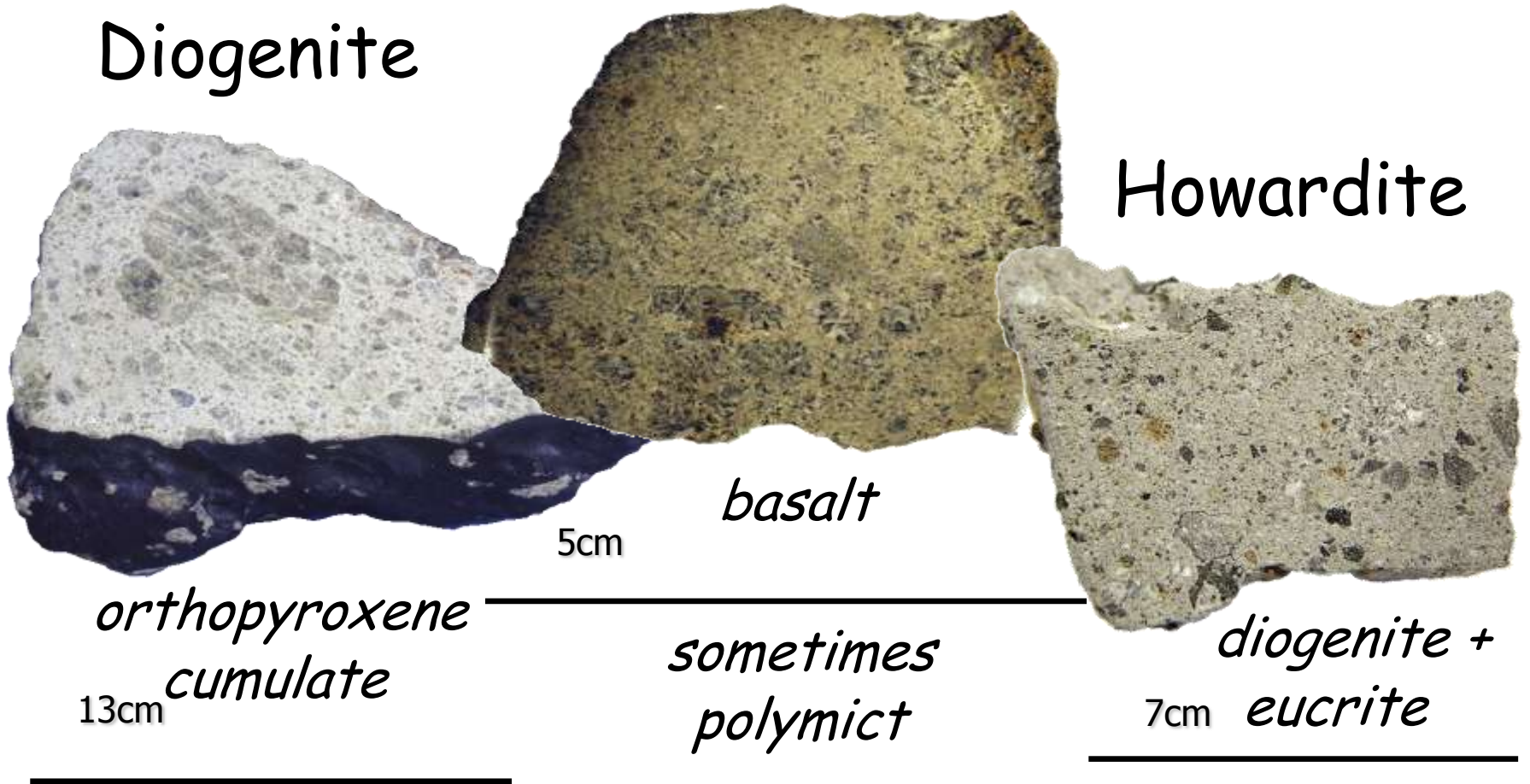
FC imaging to map surface elevations (laser altimeter was descoped for mission); topography needed to define size and shape, identify tectonic processes, and remove elevation effects from gravity signals

HED meteorites - used for instrument calibration and groundtruth for data

Eucrite

Diogenite

Howardite



*orthopyroxene
cumulate*

13cm

5cm

basalt

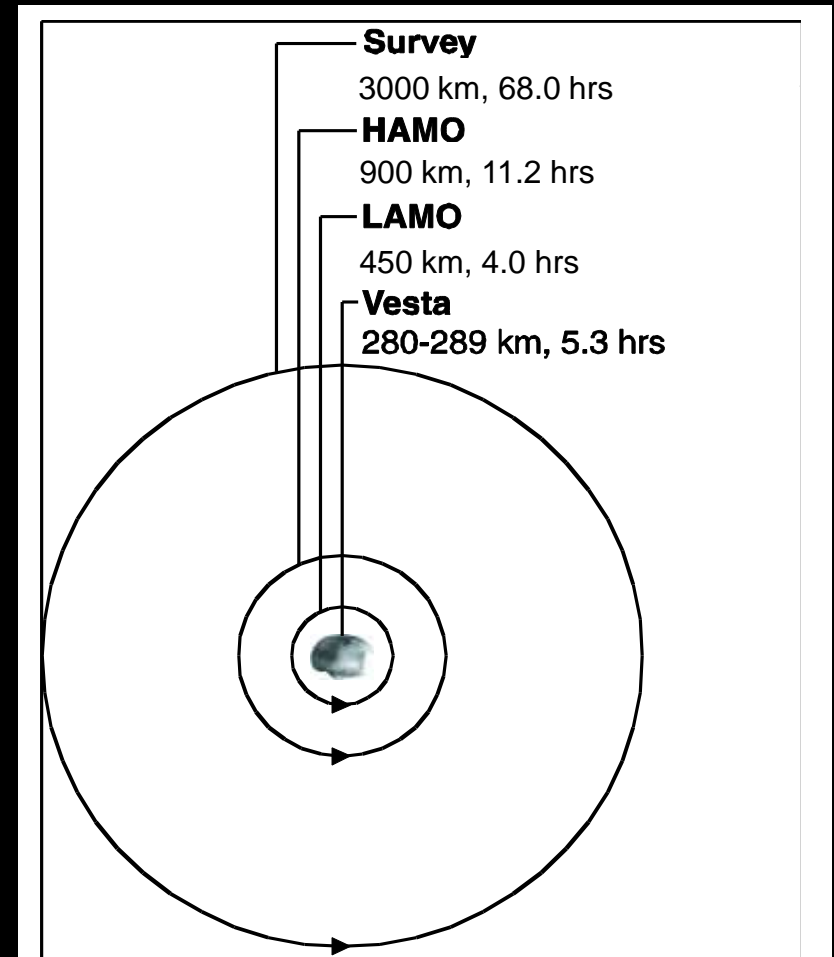
*sometimes
polymict*

7cm

*diogenite +
eucrite*

Vesta operations overview

- **Approach phase:**
 - Nav ops begin 1.4M km out, 86 days before orbit insertion
- **Survey Orbit:**
 - Duration 4 orbits (11 days)
 - VIR is prime, some FC
- **High Altitude Mapping Orbit:**
 - Duration 60 orbits (28 days)
 - FC is prime, VIR "rides along"
- **Low Altitude Mapping Orbit:**
 - Duration 420 orbits (70 days)
 - GRaND is prime, FC and VIR collect data opportunistically
 - Dedicated gravity tracking



DAWN instrument footprints

GRaND sensitivity region at LAMO

VIR slit footprints

Vesta operations

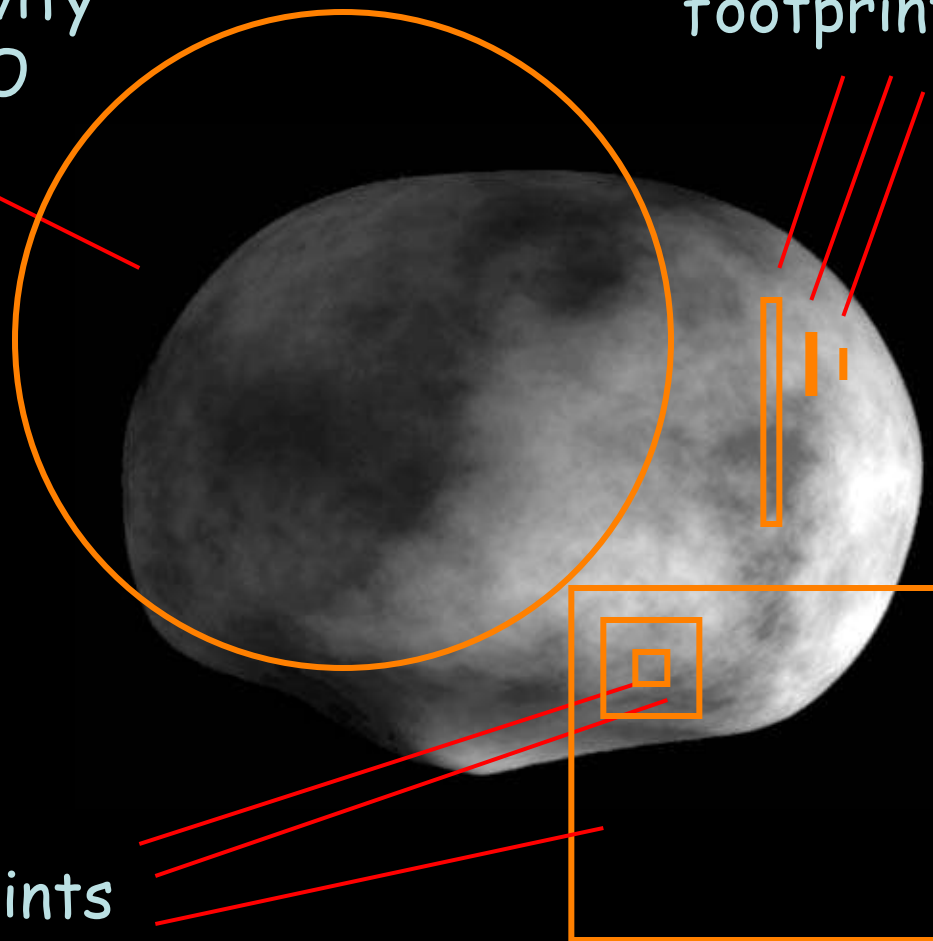
orbits - altitudes

Survey - 2720 km

HAMO - 620 km

LAMO - 170 km

FC footprints



Vesta coverage

Sunlit coverage in the funded baseline mission stops at $\sim 80^\circ$ north, allowing $\sim 65\%$ topographic and spectral coverage.

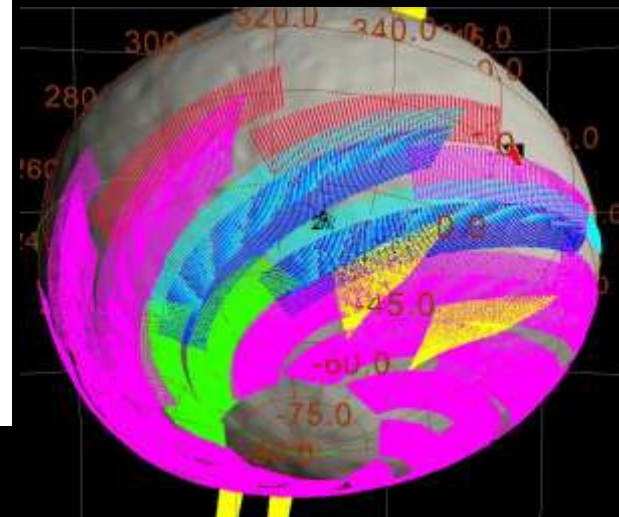
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are needed to see this picture.

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Orbital coverage at LAMO

FC at HAMO

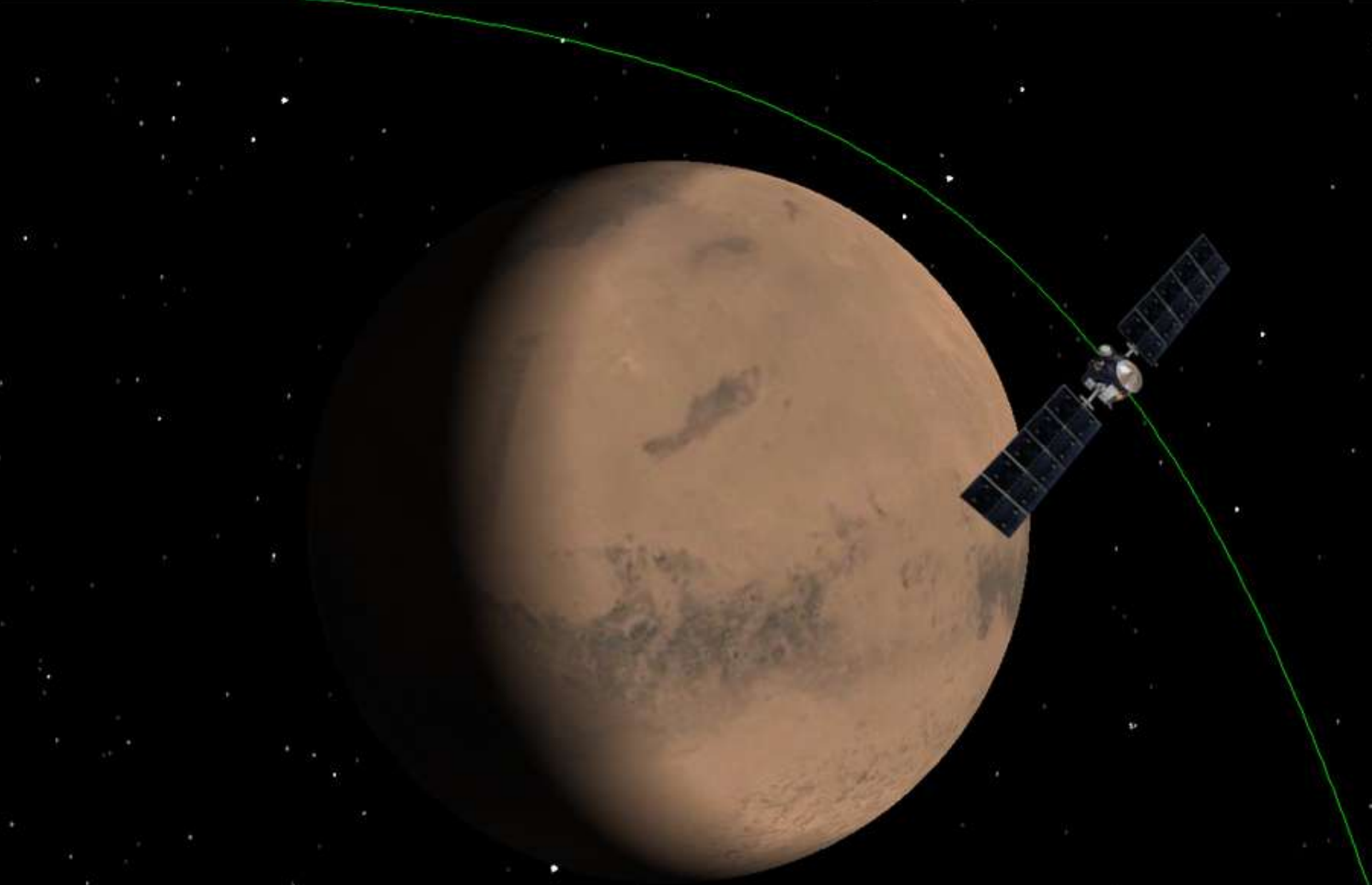


VIR at HAMO

Vesta science operations

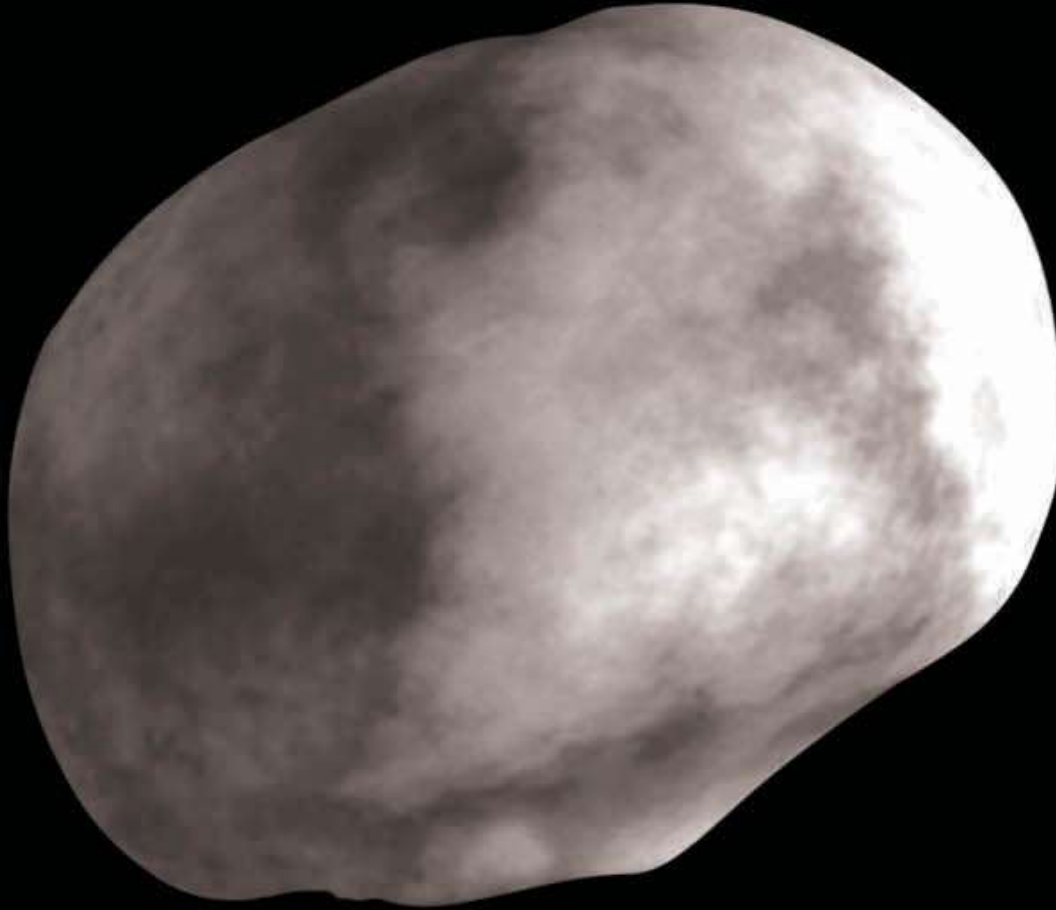
- Baseline Mission: Prelaunch approved and funded operations at Vesta from Sep 2011 to May 2012.
- Technical Baseline Mission: Inflight calibration of solar array revealed more power (increased IPS efficiency) than conservative project allotment. Thus spacecraft can travel faster, arrive at Vesta early (Aug 2011), leave late (Aug 2012), and still arrive at Ceres as planned. This mission would allow greater reserve, more low-altitude data collection, more coverage of the northern hemisphere.
- Extended Mission: By delaying arrival at Ceres by 3 to 6 months, Dawn could complete coverage of the northern hemisphere. The spacecraft was designed to operate longer than the extended mission.

NASA would need to approve additional funds if either the Technical Baseline or Extended Mission were undertaken. Community support for these options is welcome.



Mars gravity assist - February 2009

Vesta operations -
September (?) 2011-May 2012 (?)



Then on to Ceres

Hubble images of Ceres' rotation

Ceres *HST*ACS/HRC



Dec. 30, 2003 15:46UT

Ceres' layers

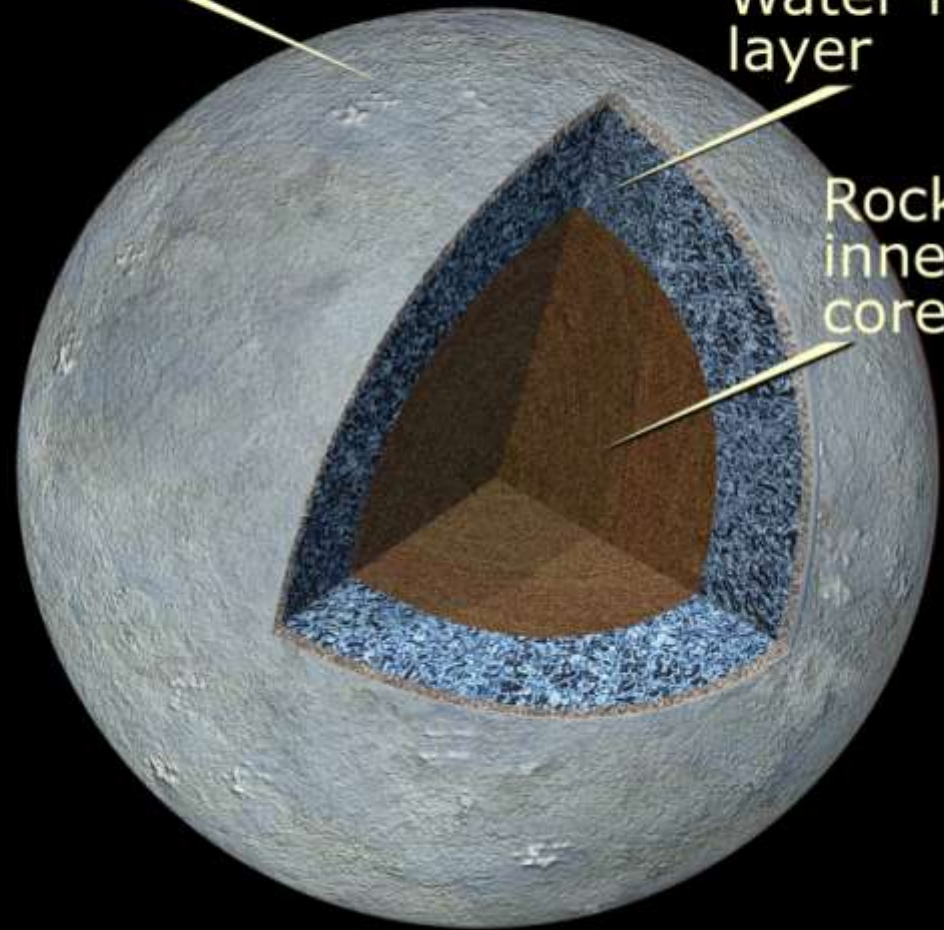
Science objectives are similar to those at Vesta, but Ceres is very different

Nominal mission
Ceres operations
commence Aug 2014

Thin, dusty
outer crust

Water-ice
layer

Rocky
inner
core



Density suggests 75% rock / 25% ice